WKB to all orders and the accuracy of the semiclassical quantization

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# WKB to all orders and the accuracy of the semiclassical quantization 

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#### Abstract

We perform a systematic WKB expansion to all orders for a one-dimensional system with potential $V(x)=U_{0} / \cos ^{2}(\alpha x)$. We are able to sum the series to the exact energy spectrum. Then we show that at any finite order the error of the WKB approximation measured in the natural units of the mean energy level spacing does not go to zero when the quantum number goes to infinity. Therefore we make the general conclusion that the semiclassical approximations fail to predict the individual energy levels within a vanishing fraction of the mean energy level spacing.


## 1. Introduction

In recent years many studies have been devoted to the transition from classical mechanics to quantum mechanics. These studies are motivated by the so-called quantum chaos (see Ozorio de Almeida 1990, Gutzwiller 1990, Casati and Chirikov 1995). An important aspect is the semiclassical quantization formula of the energy levels for integrable and quasiintegrable systems, i.e. the torus quantization initiated by Einstein (1917) and completed by Maslov (1972, 1981). As is well known, the torus quantization is just the first term of a certain $\hbar$-expansion, the so-called WKB expansion, whose higher terms can be calculated with a recursion formula at least for one degree systems (Dunham 1932, Bender et al 1977, Voros 1983).

Recently it was observed by Prosen and Robnik (1993) and also Graffi et al (1994) that the leading-order semiclassical approximation fails to predict the individual energy levels within a vanishing fraction of the mean energy level spacing. This result has been shown to be true also for the leading (torus) semiclassical approximation by Salasnich and Robnik (1996).

In this paper we analyse a simple one-dimensional system for which we are able to perform a systematic WKB expansion to all orders resulting in a convergent series whose sum is identical to the exact spectrum. For this system we show that any finite order WKB (semiclassical) approximation fails to predict the individual energy levels within a vanishing fraction of the mean energy level spacing.

[^0]
## 2. The system and the WKB expansion method

The Hamiltonian of the system is given by

$$
\begin{equation*}
H=\frac{p^{2}}{2 m}+V(x) \tag{1}
\end{equation*}
$$

where

$$
\begin{equation*}
V(x)=\frac{U_{0}}{\cos ^{2}(\alpha x)} \tag{2}
\end{equation*}
$$

Of course, the Hamiltonian is a constant of motion, whose value is equal to the total energy $E$. To perform the torus quantization it is necessary to introduce the action variable

$$
\begin{equation*}
I=\frac{1}{2 \pi} \oint p \mathrm{~d} x=\frac{\sqrt{2 m}}{\alpha}\left(\sqrt{E}-\sqrt{U_{0}}\right) . \tag{3}
\end{equation*}
$$

The Hamiltonian as a function of the action reads

$$
\begin{equation*}
H=\frac{\alpha^{2}}{2 m} I^{2}+2 \alpha \sqrt{\frac{U_{0}}{2 m}} I+U_{0} \tag{4}
\end{equation*}
$$

and after torus quantization

$$
\begin{equation*}
I=\left(v+\frac{1}{2}\right) \hbar \tag{5}
\end{equation*}
$$

where $v=0,1,2, \ldots$, the energy spectrum is given by

$$
\begin{equation*}
E_{v}^{\mathrm{tor}}=A\left[\left(v+\frac{1}{2}\right)+\frac{1}{2} B\right]^{2} \tag{6}
\end{equation*}
$$

where $A=\alpha^{2} \hbar^{2} /(2 m)$ and $B=\sqrt{8 m U_{0}} /(\alpha \hbar)$.
The Schrödinger equation of the system

$$
\begin{equation*}
\left[-\frac{\hbar^{2}}{2 m} \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}}+V(x)\right] \psi(x)=E \psi(x) \tag{7}
\end{equation*}
$$

can be solved analytically (as shown in Landau and Lifshitz 1973, Flügge 1971) and the exact energy spectrum is

$$
\begin{equation*}
E_{\nu}^{\mathrm{ex}}=A\left[\left(v+\frac{1}{2}\right)+\frac{1}{2} \sqrt{1+B^{2}}\right]^{2} \tag{8}
\end{equation*}
$$

where $v=0,1,2, \ldots$ We see that the torus quantization does not give the correct energy spectrum, but it is well known that the torus quantization is just the first term of the WKB expansion. To calculate all the terms of the WKB expansion we observe that the wavefunction can always be written as

$$
\begin{equation*}
\psi(x)=\exp \left(\frac{\mathrm{i}}{\hbar} \sigma(x)\right) \tag{9}
\end{equation*}
$$

where the phase $\sigma(x)$ is a complex function that satisfies the differential equation

$$
\begin{equation*}
\sigma^{\prime 2}(x)+\left(\frac{\hbar}{\overline{\mathrm{i}}}\right) \sigma^{\prime \prime}(x)=2 m(E-V(x)) \tag{10}
\end{equation*}
$$

The WKB expansion for the phase is given by

$$
\begin{equation*}
\sigma(x)=\sum_{k=0}^{\infty}\left(\frac{\hbar}{\mathrm{i}}\right)^{k} \sigma_{k}(x) \tag{11}
\end{equation*}
$$

Substituting equation (11) in (10) and comparing like powers of $\hbar$ gives the recursion relation ( $n>0$ )

$$
\begin{equation*}
\sigma_{0}^{\prime 2}=2 m(E-V(x)) \quad \sum_{k=0}^{n} \sigma_{k}^{\prime} \sigma_{n-k}^{\prime}+\sigma_{n-1}^{\prime \prime}=0 \tag{12}
\end{equation*}
$$

The quantization condition is obtained by requiring that the wavefunction be singlevalued:

$$
\begin{equation*}
\oint \mathrm{d} \sigma=\sum_{k=0}^{\infty}\left(\frac{\hbar}{\mathrm{i}}\right)^{k} \oint \mathrm{~d} \sigma_{k}=2 \pi \hbar v \tag{13}
\end{equation*}
$$

where $v=0,1,2, \ldots$ is the quantum number.
The zeroth-order term, which gives the Bohr-Sommerfeld formula, is given by

$$
\begin{equation*}
\oint \mathrm{d} \sigma_{0}=2 \int \mathrm{~d} x \sqrt{2 m(E-V(x))}=2 \pi \hbar\left(\sqrt{\frac{E}{A}}-\frac{1}{2} B\right) \tag{14}
\end{equation*}
$$

and the first odd term in the series gives the Maslov corrections (the Maslov index is equal to 2)

$$
\begin{equation*}
\binom{\hbar}{\mathrm{i}} \oint \mathrm{~d} \sigma_{1}=\left.\left(\frac{\hbar}{\mathrm{i}}\right) \frac{1}{4} \ln p\right|_{\text {contour }}=-\pi \hbar . \tag{15}
\end{equation*}
$$

The zeroth-and first-order terms give equation (6), which is the torus quantization formula for the energy levels (Bohr-Sommerfeld-Maslov). Here we want to analyse the quantum corrections to this formula. We observe that all the other odd terms vanish when integrated along the closed contour because they are exact differentials (Bender, Olaussen and Wang 1977). So the quantization condition (13) can be written as

$$
\begin{equation*}
\sum_{k=0}^{\infty}\left(\frac{\hbar}{\mathrm{i}}\right)^{2 k} \oint \mathrm{~d} \sigma_{2 k}=2 \pi \hbar\left(\nu+\frac{1}{2}\right) \tag{16}
\end{equation*}
$$

thus again being a sum over even-numbered terms only. The next two non-zero terms (Narimanov 1995, Bender et al 1977, Robnik and Salasnich 1996) are
$\left(\frac{\hbar}{\mathrm{i}}\right)^{2} \oint \mathrm{~d} \sigma_{2}=-\frac{\hbar^{2}}{\sqrt{2 m}} \frac{1}{12} \frac{\partial^{2}}{\partial E^{2}} \int \mathrm{~d} x \frac{V^{\prime 2}(x)}{\sqrt{E-V(x)}}$
$\left(\frac{\hbar}{\mathrm{i}}\right)^{4} \oint \mathrm{~d} \sigma_{4}=\frac{\hbar^{4}}{(2 m)^{3 / 2}}\left[\frac{1}{120} \frac{\partial^{3}}{\partial E^{3}} \int \mathrm{~d} x \frac{V^{\prime \prime 2}(x)}{\sqrt{E-V(x)}}-\frac{1}{288} \frac{\partial^{4}}{\partial E^{4}} \int \mathrm{~d} x \frac{V^{\prime 2}(x) V^{\prime \prime}(x)}{\sqrt{E-V(x)}}\right]$.

A straightforward calculation of these terms gives (see the appendix)

$$
\begin{equation*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{2} \oint \mathrm{~d} \sigma_{2}=-\frac{2 \pi \hbar}{4 B} \tag{19}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{4} \oint \mathrm{~d} \sigma_{4}=\frac{2 \pi \hbar}{16 B^{3}} . \tag{20}
\end{equation*}
$$

Up to fourth order in $\hbar \sim B^{-1}$ the quantization condition reads

$$
\begin{equation*}
E_{v}^{(4)}=A\left[\left(v+\frac{1}{2}\right)+\frac{1}{2} B+\frac{1}{4 B}-\frac{1}{16 B^{3}}\right]^{2} . \tag{21}
\end{equation*}
$$

The first two terms on the right-hand side give the torus quantization formula, and the other two terms are quantum corrections. Higher-order quantum corrections quickly increase in complexity but in this specific case they can be calculated. We first verify by induction, following Bender et al (1977), that the solution to (12) has the general form

$$
\begin{equation*}
\sigma_{n}^{\prime}(x)=\left(\sigma_{0}^{\prime}\right)^{1-3 n} P_{n}(\cos (\alpha x)) \sin ^{f(n)}(\alpha x) \tag{22}
\end{equation*}
$$

where $f(n)=0$ for $n$ even and $f(n)=1$ for $n$ odd, and $P_{n}$ is a polynomial given by

$$
\begin{equation*}
P_{n}(\cos (\alpha x))=\sum_{l=0}^{g(n)} C_{n, l} \cos ^{2 l-3 n}(\alpha x) \tag{23}
\end{equation*}
$$

with $g(n)=(3 n-2) / 2$ for $n$ even and $g(n)=(3 n-3) / 2$ for $n$ odd.
The integrals in (16) are performed by substituting $z=\tan (\alpha x)$. In this way the $2 k$-term reduces to

$$
\begin{equation*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{2 k} \oint \mathrm{~d} \sigma_{2 k}=\left(\frac{\hbar}{\mathrm{i}}\right)^{2 k} \frac{(2 m)^{1 / 2-3 k}}{\alpha} \sum_{l=0}^{3 k-1} C_{2 k, l} \oint \mathrm{~d} z \frac{\left(1+z^{2}\right)^{3 k-l-1}}{\left(E-U_{0}-U_{0} z^{2}\right)^{3 k-1 / 2}} . \tag{24}
\end{equation*}
$$

We observe that

$$
\begin{align*}
& \oint \mathrm{d} z \frac{\left(1+z^{2}\right)^{3 k-l-1}}{\left(E-U_{0}-U_{0} z^{2}\right)^{3 k-1 / 2}} \\
& \quad=(-1)^{3 k-1} \frac{\Gamma\left(\frac{1}{2}\right)}{\Gamma\left(3 k-\frac{1}{2}\right)} \frac{\partial^{3 k-1}}{\partial E^{3 k-1}} \oint \mathrm{~d} z \frac{\left(1+z^{2}\right)^{3 k-l-1}}{\left(E-U_{0}-U_{0} z^{2}\right)^{1 / 2}} \tag{25}
\end{align*}
$$

so the only non-zero term is for $l=0$ :

$$
\begin{gather*}
\frac{\partial^{3 k-1}}{\partial E^{3 k-1}} \oint \mathrm{~d} z \frac{\left(1+z^{2}\right)^{3 k-1}}{\left(E-U_{0}-U_{0} z^{2}\right)^{1 / 2}}=\frac{2^{6 k-1}}{U_{0}^{1 / 2}} \frac{\Gamma(3 k-1 / 2)^{2}}{\Gamma(6 k-1)} \frac{\partial^{3 k-1}}{\partial E^{3 k-1}} \beta^{3 k-1} \\
\quad=\frac{2^{6 k-1}}{U_{0}^{1 / 2}} \frac{\Gamma(3 k-1 / 2)^{2}}{\Gamma(6 k-1)} \Gamma(3 k) \frac{1}{U_{0}^{3 k-1 / 2}} 2 \pi \tag{26}
\end{gather*}
$$

where $\beta=\left(E-U_{0}\right) / U_{0}$. At this stage we obtain

$$
\begin{equation*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{2 k} \oint \mathrm{~d} \sigma_{2 k}=(-1)^{5 k-1} \hbar^{2 k} \frac{(2 m)^{1 / 2-3 k}}{\alpha} C_{2 k, 0} \frac{1}{U_{0}^{3 k-1 / 2}} 2 \pi . \tag{27}
\end{equation*}
$$

Now we need to find the coefficient $C_{2 k, 0}$ explicitly. By inserting (22) with (23) in the recursion relation (12) we obtain
$\sum_{k=0}^{n} C_{k, 0} C_{n-k, 0}-\left(2 m U_{0} \alpha\right) C_{n-1,0}=\sum_{k=1}^{n-1} C_{k, 0} C_{n-k, 0}+2 C_{n, 0}-\left(2 m U_{0} \alpha\right) C_{n-1,0}=0$
from which we have

$$
\begin{equation*}
C_{k, 0}=\frac{1}{2}\left[\left(2 m \alpha U_{0}\right) C_{k-1,0}-\sum_{j=1}^{k-1} C_{j, 0} C_{k-j, 0}\right] \quad C_{0,0}=1 \tag{29}
\end{equation*}
$$

From this equation one shows that $C_{1,0}=m \alpha U_{0}$. Furthermore, it easy to show that all higher odd coefficients vanish: $C_{2 k+3,0}=0$ for $k=0,1,2, \ldots$. The solution of this equation for the remaining non-zero even coefficients is given by

$$
\begin{equation*}
C_{2 k, 0}=(-1)^{k}\left(2 m U_{0} \alpha\right)^{2 k} 2^{-2 k}\binom{\frac{1}{2}}{k} \tag{30}
\end{equation*}
$$

which can be verified by direct substitution in equation (29), resulting in an identity for half-integer binomial coefficients. Then the integral (27) can be written as

$$
\begin{equation*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{2 k} \oint \mathrm{~d} \sigma_{2 k}=(-1) \hbar^{2 k} 2 \pi \alpha^{2 k-1}(2 m)^{1 / 2-k} 2^{-2 k}\binom{\frac{1}{2}}{k} U_{0}^{k-1 / 2}=-\frac{1}{2}\binom{\frac{1}{2}}{k} \frac{2 \pi \hbar}{B^{2 k-1}} . \tag{31}
\end{equation*}
$$

In conclusion, the WKB quantization to all orders (16) is

$$
\begin{equation*}
E_{v}^{(\infty)}=A\left[\left(v+\frac{1}{2}\right)+\frac{1}{2} \sum_{k=0}^{\infty}\binom{\frac{1}{2}}{k} \frac{1}{B^{2 k-1}}\right]^{2} . \tag{32}
\end{equation*}
$$

Because $\sum_{k=0}^{\infty}\left(\frac{1}{2}\right) B^{1-2 k}=\sqrt{1+B^{2}}$ we have $E_{v}^{\text {ex }}=E_{v}^{(\infty)}$, i.e. the WKB series converges to the exact result (8).

Now we can calculate the error in units of the mean level spacing $\Delta E_{v}=E_{v+1}^{\mathrm{ex}}-E_{v}^{\mathrm{ex}}$ between the exact level $E_{v}^{\mathrm{ex}}$ and its WKB approximation $E_{v}^{(N)}$ to $N$ th order:

$$
\begin{equation*}
\frac{E_{v}^{\mathrm{ex}}-E_{v}^{(N)}}{\Delta E_{v}}=\frac{1}{2} \sum_{k=N+1}^{\infty}\binom{\frac{1}{2}}{k} \frac{1}{B^{2 k-1}} \quad \text { for } v \rightarrow \infty \tag{33}
\end{equation*}
$$

The limit clearly shows that even for arbitrarily small but finite $\hbar(1 \ll B<\infty)$, the relative error for any finite WKB approximation becomes constant on increasing $v$, and scales as

$$
\begin{equation*}
\frac{E_{v}^{\mathrm{ex}}-E_{v}^{(N)}}{\Delta E_{v}} \sim \frac{1}{2}\binom{\frac{1}{2}}{N+1} \frac{1}{B^{2 N+1}} \quad B \rightarrow \infty \tag{34}
\end{equation*}
$$

Note that the limit $B \rightarrow \infty$ is equivalent to the limit $\hbar \rightarrow 0$.

## 3. Conclusions and discussion

For our present system we can conclude that to any finite order semiclassical approximation the error measured in units of the mean level spacing remains constant even if the quantum number increases indefinitely, contrary to the naive expectation, even if $\hbar$ is small but fixed and non-zero. This confirms the general statements made by Prosen and Robnik (1993). We have thus provided a clear demonstration that the semiclassical methods cannot predict the individual energy levels (and also their wavefunctions) within a vanishing fraction of the mean energy level spacing in the limit when the quantum number goes to infinity, $v \rightarrow \infty$, but $\hbar$ is small and fixed. Therefore we cannot expect the semiclassics at any non-zero value of $\hbar$ to correctly describe the fine structure of energy spectra manifested in the short-range statistics like the energy level repulsion, which was predicted to be a purely quantum effect (Robnik 1986), later reconfirmed by Berry (1991). On the other hand, Prosen and Robnik (1993) have shown that the long-range statistics of the energy spectra are very well captured even by the lowest-order semiclassical approximation. This is of course compatible with the very important semiclassical theory of delta statistics $\Delta(L)$ (spectral rigidity) by Berry (1985), employing the Gutzwiller periodic orbit theory (Gutzwiller 1990), where agreement with predictions of random matrix theories and with the experimental and numerical data has been obtained at large L. Also, Berry and Tabor (1977) have used torus quantization of integrable systems (with more than one degree of freedom), predicting the Poissonian (exponential) energy level spacing distribution. Our results show that their result cannot be rigorous, especially since we know some counterexamples of integrable systems with non-Poissonian statistics (Bleher et al 1993), and also know that their approximation does not take into account the non-perturbative tunnelling effects, but it is nevertheless a good heuristic argument explaining why typically we do observe Poissonian statistics in classically
integrable systems. By typically we mean that the set of exceptions has a small or even vanishing measure.

We should emphasize again that the error of the torus quantization (which Berry and Tabor (1977) have used) and of the Gutzwiller trace formula (which obviously has generally the same quality as EBK quantization of classically integrable systems) scales as $\propto \hbar^{2}$, in the limit of fixed quantum numbers but $\hbar \rightarrow 0$. (This is the actual semiclassical limit in the strong sense.) In systems with $f$ degrees of freedom the mean energy level spacing $\Delta E$ scales as $\hbar^{f}$ (due to the asymptotically exact Thomas-Fermi rule). Therefore, the error of the leading semiclassical approximation (EBK, Gutzwiller) measured in units of $\Delta E$ scales as $\hbar^{(2-f)}$. Thus it goes to zero only in systems with one degree of freedom $(f=1)$, it is constant for two degrees of freedom $(f=2)$, and can even diverge for higher degrees of freedom $(f \geqslant 3)$. Thus to improve the theoretical results we need higher corrections to the leading semiclassical terms. However, one should not forget that in order to explain quantal results $\hbar$ must be non-zero and fixed, however small, in which case the fine structure of energy spectra at scales smaller than certain characteristic range (a power of $\hbar$ ) cannot be correctly described.

The conclusion of this paper is that the semiclassical methods are just not good enough (at any order) to describe the fine structure of energy spectra and wavefunctions. Our approach leading to the above conclusion rests upon a systematic WKB expansion for the potential $V(x)=U_{0} / \cos ^{2}(\alpha x)$ using the technique of Bender et al (1977). We are able to calculate all orders, the series is convergent and can be summed precisely to the exact result.

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## Appendix

In this appendix we show how to obtain equations (19) and (20). In all integrals of this section the limits of integration are between the two turning points. After the substitution $z=\tan (\alpha x)$, we have

$$
\begin{align*}
\int \mathrm{d} x \frac{V^{\prime 2}(x)}{\sqrt{E-V(x)}} & =\frac{4 \alpha U_{0}^{2}}{\sqrt{U_{0}}} \int_{-\sqrt{\beta}}^{\sqrt{\beta}} \mathrm{d} z \frac{z^{2}\left(z^{2}+1\right)}{\sqrt{\beta-z^{2}}} \\
& =\frac{4 \alpha U_{0}^{2}}{\sqrt{U_{0}}}\left(3 \beta^{2}+4 \beta\right) \frac{\pi}{8} \tag{A1}
\end{align*}
$$

where $\beta=\left(E-U_{0}\right) / U_{0}$. In conclusion we have

$$
\begin{equation*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{2} \oint \mathrm{~d} \sigma_{2}=-\frac{\hbar^{2} \alpha \pi}{8 \sqrt{2 m U_{0}}}=-\frac{2 \pi \hbar}{4 B} \tag{A2}
\end{equation*}
$$

with $B=\sqrt{8 m U_{0}} /(\alpha \hbar)$.

To obtain equation (21) we proceed in the same way:

$$
\begin{align*}
\int \mathrm{d} x \frac{V^{\prime \prime 2}(x)}{\sqrt{E-V(x)}} & =\frac{4 \alpha^{3} U_{0}^{2}}{\sqrt{U_{0}}} \int_{-\sqrt{\beta}}^{\sqrt{\beta}} \mathrm{d} z \frac{\left(9 z^{4}+6 z^{2}+1\right)\left(z^{2}+1\right)}{\sqrt{\beta-z^{2}}} \\
& =\frac{4 \alpha U_{0}^{2}}{\sqrt{U_{0}}}\left(45 \beta^{3}+90 \beta^{2}+56 \beta+16\right) \frac{\pi}{16} \tag{A3}
\end{align*}
$$

from which we obtain

$$
\begin{equation*}
\frac{\partial^{3}}{\partial E^{3}} \int \mathrm{~d} x \frac{V^{\prime \prime 2}(x)}{\sqrt{E-V(x)}}=\frac{135 \pi \alpha^{3} \sqrt{U_{0}}}{2 U_{0}^{2}} \tag{A4}
\end{equation*}
$$

For the last integral we have

$$
\begin{align*}
\int \mathrm{d} x \frac{V^{\prime 2}(x) V^{\prime \prime}(x)}{\sqrt{E-V(x)}} & =\frac{8 \alpha^{3} U_{0}^{2}}{\sqrt{U_{0}}} \int_{-\sqrt{\beta}}^{\sqrt{\beta}} \mathrm{d} z \frac{z^{2}\left(3 z^{2}+1\right)\left(z^{2}+1\right)^{2}}{\sqrt{\beta-z^{2}}} \\
& =\frac{8 \alpha U_{0}^{2}}{\sqrt{U_{0}}}\left(105 \beta^{4}+280 \beta^{3}+240 \beta^{2}+60 \beta\right) \frac{\pi}{128} \tag{A5}
\end{align*}
$$

from which we obtain

$$
\begin{equation*}
\frac{\partial^{4}}{\partial E^{4}} \int \mathrm{~d} x \frac{V^{\prime 2}(x) V^{\prime \prime}(x)}{\sqrt{E-V(x)}}=\frac{315 \pi \alpha^{3} \sqrt{U_{0}}}{2 U_{0}^{2}} \tag{A6}
\end{equation*}
$$

In conclusion we have

$$
\begin{align*}
\left(\frac{\hbar}{\mathrm{i}}\right)^{4} \oint \mathrm{~d} \sigma_{4} & =\frac{\hbar^{4}}{(2 m)^{3 / 2}} \frac{\alpha^{3} \sqrt{U_{0}}}{U_{0}^{2}}\left[\frac{1}{120} \frac{135}{2}-\frac{1}{288} \frac{315}{2}\right] \\
& =\frac{\hbar^{4} \alpha^{3} \pi \sqrt{U_{0}}}{64(2 m)^{3 / 2} U_{0}^{2}}=\frac{2 \pi \hbar}{16 B^{3}} . \tag{A7}
\end{align*}
$$

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